

# Fermionic zero modes in baby Skyrme models

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- 1 Introduction
  - Motivation
- 2 The model
  - without SUSY
  - with SUSY
- 3 Fermionic sector
  - The full lagrangian
  - SUSY transformations
  - Counting degrees
  - zero modes
  - example
- 4 Exotic models
  - Adding a D-term perturbation
- 5 Full SUSY baby
- 6 Conclusions

# Motivation

- The baby Skyrme model is a Skyrme-type model in  $2 + 1$  dimensions.
- Static solutions are classified by a winding number.
- Stability of solutions requires a potential term.
- $\mathcal{L}_{bS} = \mathcal{L}_0 + \mathcal{L}_2 + \mathcal{L}_4$ .
- It has applications from condensed matter to cosmology.
- SUSY version of the model  $\rightarrow$  canonical way of introducing fermions
- We can relate SUSY transformations with the BPS structure of the model (BPS restriction).
- Our goal: how to get some information about the fermionic zero modes in the SUSY model?

# The baby Skyrme model

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## Baby Skyrme Lagrangian

$$\mathcal{L} = \frac{\lambda_2}{4} \partial_\mu \vec{\phi} \partial^\mu \vec{\phi} - \frac{\lambda_4}{8} \left( \partial_\mu \vec{\phi} \times \partial_\nu \vec{\phi} \right)^2 - \lambda_0 \mathcal{V}(\phi^3)$$

$$\vec{\phi} \in \mathbb{S}^2 \quad \vec{\phi}_0 : \mathbb{S}^2 \rightarrow \mathbb{S}^2$$

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After stereographic projection and  $\lambda_2 = 0$  BPS submodel  $\rightarrow$

$$\mathcal{L}_{BBS} = -\frac{1}{(1+u\bar{u})^4} \left( (\partial_\mu u \partial^\mu \bar{u})^2 - (\partial_\mu u)^2 (\partial_\nu \bar{u})^2 \right) - \mathcal{V}(u\bar{u})$$

Energy in terms of holomorphic/antiholomorphic derivatives

$$E = \frac{1}{4} \int d^2x \left( \frac{(|\partial u|^2 - |\bar{\partial} u|^2)^2}{(1+u\bar{u})^4} + 4\mathcal{V}(u\bar{u}) \right)$$

where  $\partial = \partial_1 + i\partial_2$ . Therefore...

$$E \geq \mp 4\pi \int d^2x \sqrt{V(u, \bar{u})} Q, \quad Q = \frac{1}{4\pi} \frac{(|\partial u|^2 - |\bar{\partial} u|^2)}{(1 + u\bar{u})^2}$$

$Q$  is the degree of  $u$ . The energy bound is saturated if

$$\partial u = e^{i\eta} \sqrt{\sqrt{V(u, \bar{u})} (1 + u\bar{u})^2 + \frac{1}{2} (|\partial u|^2 + |\bar{\partial} u|^2)}$$

and  $\eta$  a free phase ( $\mathcal{L}$  invariant under  $u \rightarrow e^{i\eta} u$ ). This is the BPS equation of the model (keep in mind!). Let's move to SUSY BPS baby.

# The SUSY baby Skyrme model

$$\mathcal{L}_{BbS}^{\text{SUSY}} = \int d^4\theta K(\Phi, \Phi^\dagger) + \int d^4\theta H(\Phi, \Phi^\dagger) D^\alpha \Phi D_\alpha \Phi \bar{D}^{\dot{\beta}} \Phi^\dagger \bar{D}_{\dot{\beta}} \Phi^\dagger$$

$$\Phi = u + i\theta\sigma^\mu\bar{\theta}\partial_\mu u + \frac{1}{4}\theta\theta\bar{\theta}\bar{\theta}\Box u + \sqrt{2}\theta\psi - \frac{i}{\sqrt{2}}\theta\theta\partial_\mu\psi\sigma^\mu\bar{\theta} + \theta\theta F$$

- Integrate w.r.t Grassmann coordinates
- Switch off fermions

# The SUSY baby Skyrme model

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- Integrate w.r.t Grassmann coordinates
- Switch off fermions

$$\begin{aligned} \mathcal{L}_{BbS}^{\text{SUSY}} &= g(u, \bar{u}) (\partial^\mu u \partial_\mu \bar{u} + F \bar{F}) + \\ &+ h(u, \bar{u}) \left( (\partial_\mu u)^2 (\partial_\nu \bar{u})^2 + 2F \bar{F} \partial^\mu \bar{u} \partial_\mu u + (F \bar{F})^2 \right) \end{aligned}$$

- Solve eom's for  $F$

- $F = 0$
- $F = e^{i\eta} \sqrt{-\partial_\mu u \partial^\mu \bar{u} - \frac{g(u, \bar{u})}{2h(u, \bar{u})}}$

Take the non-zero solution and substitute (with the choice  $h = 1/(1 + u\bar{u})^4$ ) Finally:

$$\mathcal{L}_{BbS}^{\text{SUSY}} = h(u, \bar{u}) \left( (\partial_\mu u)^2 (\partial_\nu \bar{u})^2 - (\partial^\mu \bar{u} \partial_\mu u)^2 \right) - \frac{g(u, \bar{u})^2}{4h(u, \bar{u})}$$

- No  $\sigma$ -model term.
- The potential term comes from the Kähler metric  $g$ .
- No  $N = 2$  baby model (with quadratic term)?
- The BPS equation in the SUSY model is simply

$$\partial u = F$$

The full lagrangian

# Fermions

Taking into account the fermions we get

$$\begin{aligned}\mathcal{L}_2 = \int d^4\theta K(\Phi, \Phi^\dagger) &= g(u, \bar{u}) \left( \partial^\mu u \partial_\mu \bar{u} - \frac{i}{2} \psi \sigma^\mu \mathcal{D}^\mu \bar{\psi} \right. \\ &\quad \left. + \frac{i}{2} \mathcal{D}^\mu \psi \sigma_\mu \bar{\psi} + F \bar{F} \right) + \mathcal{O}(\psi \bar{\psi})^2\end{aligned}$$

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$$\begin{aligned} \mathcal{L}_4^0 = & \int d^4\theta H(\Phi, \Phi^\dagger) D^\alpha \Phi D_\alpha \Phi \bar{D}^{\dot{\beta}} \Phi^\dagger \bar{D}_{\dot{\beta}} \Phi^\dagger|_{0-\text{der}} = \\ & h(u, \bar{u}) ((\partial_\mu u)^2 (\partial_\nu \bar{u})^2 + 2F \bar{F} \partial^\mu u \partial_\nu \bar{u} + (F \bar{F})^2) \\ & + \frac{i}{2} \psi \sigma^\mu \bar{\psi} (\bar{u}_{,\mu} \square u - u_{,\mu} \square \bar{u}) \\ & - \frac{i}{2} \psi \sigma^\mu \bar{\psi}_{,\nu} u_{,\mu} \bar{u}_{,\nu} - i \psi^{\mu} \sigma^\nu \bar{\psi} u_{,\mu} \bar{u}_{,\nu} - \frac{i}{2} \psi \sigma^\mu \bar{\sigma}^\rho \sigma^\nu \bar{\psi}_{,\nu} u_{,\mu} \bar{u}_{,\rho} + \dots \end{aligned}$$

## The full lagrangian

$$\begin{aligned}
 & \dots + F \square u \bar{\psi} \bar{\psi} + \frac{1}{2} Fu_{,\mu} \partial^\mu (\bar{\psi} \bar{\psi}) + \bar{F} \square \bar{u} \psi \psi + \frac{1}{2} \bar{F} \bar{u}_{,\mu} \partial^\mu (\psi \psi) \\
 & + \frac{1}{2} Fu_{,\mu} (\psi \bar{\sigma}^\mu \sigma^\nu \bar{\psi}_{,\nu} - \bar{\psi}_{,\nu} \bar{\sigma}^\mu \sigma^\nu \bar{\psi}) + \frac{1}{2} \bar{F} \bar{u}_{,\mu} (\psi_{,\nu} \sigma^\nu \bar{\sigma}^\mu \psi - \psi \sigma^\nu \bar{\sigma}^\mu \psi_{,\nu}) \\
 & + \frac{3i}{2} F \bar{F} (\psi_{,\mu} \sigma^\mu \bar{\psi} - \psi \sigma^\mu \bar{\psi}_{,\mu}) + \frac{i}{2} \psi \sigma^\mu \bar{\psi} (F \bar{F}_{,\mu} - \bar{F} F_{,\mu}) \\
 & . + \frac{i}{2} \psi_{,\nu} \sigma^\nu \bar{\sigma}^\mu \sigma^\rho u_{,\mu} \bar{u}_{,\rho} + \mathcal{O}(\psi)^3
 \end{aligned}$$

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& + \frac{3i}{2} F \bar{F} (\psi_{,\mu} \sigma^\mu \bar{\psi} - \psi \sigma^\mu \bar{\psi}_{,\mu}) + \frac{i}{2} \psi \sigma^\mu \bar{\psi} (F \bar{F}_{,\mu} - \bar{F} F_{,\mu}) \\
& . + \frac{i}{2} \psi_{,\nu} \sigma^\nu \bar{\sigma}^\mu \sigma^\rho u_{,\mu} \bar{u}_{,\rho} + \mathcal{O}(\psi)^3)
\end{aligned}$$

$$\begin{aligned}
\mathcal{L}_4^1 &= \partial_u h(u, \bar{u}) \left( \frac{1}{2} F \bar{F}^2 \psi^2 - \frac{1}{2} F \psi^2 (\bar{u}_{,\nu})^2 + i \sigma^\mu \bar{\psi} \partial_\mu \psi \right. \\
&\quad - i \sqrt{2} (\bar{u}_{,\mu})^2 \bar{\sigma}^\nu \psi^2 u_{,\nu} \\
&\quad - 2i F \bar{F} \bar{u}_{,\mu} \psi \sigma^\mu \psi - 2 F \bar{u}_{,\mu} \psi \bar{\psi} \bar{\sigma}^\mu \sigma_\nu u_{,\nu} + 2i \psi \bar{\psi} \sigma^\mu \sigma^\nu \sigma^\rho u_{,\mu} \bar{u}_{,\nu} u_{,\rho} \\
&\quad \left. + 2 F \bar{F} \psi \sigma^\mu \bar{\psi} u_{,\mu} - F \bar{\psi}^2 \sigma^\mu \sigma^\nu u_{,\mu} u_{,\mu} + \bar{F} \psi^2 \sigma^\mu \sigma^\nu \bar{u}_{,\mu} u_{,\nu} + \mathcal{O}(\psi)^3 \right)
\end{aligned}$$

## The full lagrangian

+ 19 terms more.

- The auxiliary field  $F$  becomes dynamical.
- No trivial solution for  $F$  (due to linear terms).
- Even at quadratic order in fermions it is difficult to obtain information from eom's.
- We can use SUSY transformations to obtain fermionic zero modes from bosonic solutions.

$$\delta\psi^\alpha = \partial^{\alpha\dot{\alpha}} u \bar{\xi}_{\dot{\alpha}} + F \xi^\alpha$$

Preservation of (a fraction of SUSY)  $\Rightarrow \delta\psi^\alpha|_{\text{on-shell}} = 0$ . In terms  
holomorphic/antiholomorphic derivatives

$$\delta\psi^\alpha = \rho^\alpha \partial u + \eta^\alpha \bar{\partial} u + (e^{i\eta} \partial u + F) \xi^\alpha$$

where

$$\begin{aligned}\eta^\alpha &= -\frac{1}{2} \left( \sigma_1^{\alpha\dot{\beta}} - i\sigma_3^{\alpha\dot{\beta}} \right) \bar{\xi}_{\dot{\beta}} \\ \rho^\alpha &= +\frac{1}{2} \left( \sigma_1^{\alpha\dot{\beta}} + i\sigma_3^{\alpha\dot{\beta}} \right) \bar{\xi}_{\dot{\beta}} + e^{i\eta'} \xi^\alpha.\end{aligned}$$

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Therefore

$$\delta\psi^\alpha|_{\text{on-shell}} = \rho^\alpha \partial u + \eta^\alpha \bar{\partial} u$$

## Counting degrees

- If  $\partial u \neq e^{i\eta} F$ ,  $\delta\psi^\alpha|_{\text{on-shell}} = 0 \Rightarrow \rho^\alpha = \eta^\alpha = \xi^\alpha = 0$  and SUSY is not generically preserved.
- If  $\partial u = e^{i\eta} F$ ,  $\delta\psi^\alpha|_{\text{on-shell}} = 0 \Rightarrow \rho^\alpha = \eta^\alpha = 0$ ,  $\eta^\alpha$  has 1 d.o.f. and  $\rho^\alpha$  has 2 d.o.f.  $\Rightarrow$  3 SUSY generators are broken  $\Rightarrow$  solutions are 1/4 SUSY (equiv.  $\dim \text{Ker}\delta = 1$ ).
- 3 generators broken  $\Rightarrow$  3 goldstone fermions (fermionic zero modes).

In the background of the BPS solutions we have obtained:

$\delta\psi^\alpha|_{\text{on-shell}} = \rho^\alpha \partial u + \eta^\alpha \bar{\partial} u$  and 3 Grassmann parameters leads to three fermionic d.o.f.

$$\left\{ \frac{1}{2}\chi \begin{pmatrix} 1 \\ i \end{pmatrix} \bar{\partial} u, \frac{1}{2} \begin{pmatrix} \zeta_1 \\ 0 \end{pmatrix} \partial u, \frac{1}{2} \begin{pmatrix} 0 \\ \zeta_2 \end{pmatrix} \partial u \right\}$$

In the symmetric *ansatz*  $u = f(r)e^{in\varphi}$  with boundary conditions

$$f(r=0) = \infty, \quad f(r=R) = 0 \quad \text{and} \quad f'(r=R) = 0$$

$R = \text{finite}, \infty$  for a compact or non-compact solution respect. The fermionic zero modes can be written as

$$\begin{pmatrix} \frac{1}{2}\chi \begin{pmatrix} 1 \\ i \end{pmatrix} \frac{e^{i(n-1)\theta}(rf'(r)+nf(r))}{r} \\ \frac{1}{2}\zeta_1 \begin{pmatrix} 1 \\ 0 \end{pmatrix} \frac{e^{i(n+1)\varphi}(rf'(r)-nf(r))}{r} \\ \frac{1}{2}\zeta_2 \begin{pmatrix} 0 \\ 1 \end{pmatrix} \frac{e^{i(n+1)\varphi}(rf'(r)-nf(r))}{r} \end{pmatrix}$$

- If  $R = \infty$  (non compact solution),  $\delta\psi^\alpha|_{\text{BPS}} = 0 \Leftrightarrow \rho^\alpha = \eta^\alpha = 0, 1/4$  SUSY in  $\mathbb{R}^2$ .
- if  $R = \text{finite}$ , we have  $f'(r) = f(r) = 0$  for  $r \geq R \Rightarrow 1/4$  SUSY inside the compacton, but 1 SUSY outside.
  - $\Rightarrow$  fermionic zero modes only exist inside the compacton

Let us take the following potential:

$$\mathcal{V}(u, \bar{u}) = \left( \frac{u\bar{u}}{1+u\bar{u}} \right)^s, \quad s \in (1, 2)$$

The corresponding Kähler potential in the SUSY model is given by

$$K(\Phi, \Phi^\dagger) = 8 \frac{(\Phi\Phi^\dagger)^{\frac{s+2}{2}}}{(s+2)^2} {}_2F_1 \left( \frac{s+2}{2}, \frac{s+2}{2}, \frac{s+4}{2}, -\Phi^\dagger\Phi \right).$$

For the  $n = 1$  charge solution we get the following profile function:

$$f(r) = \begin{cases} \frac{\left(1 - \frac{r^2}{R^2}\right)^{\frac{1}{2-s}}}{\sqrt{1 - \left(1 - \frac{r^2}{R^2}\right)^{-\frac{2}{s-2}}}}, & 0 \leq r \leq R \\ 0, & r \geq R \end{cases}$$

Compacton radius  $R^2 = 4/(2-s)$ .

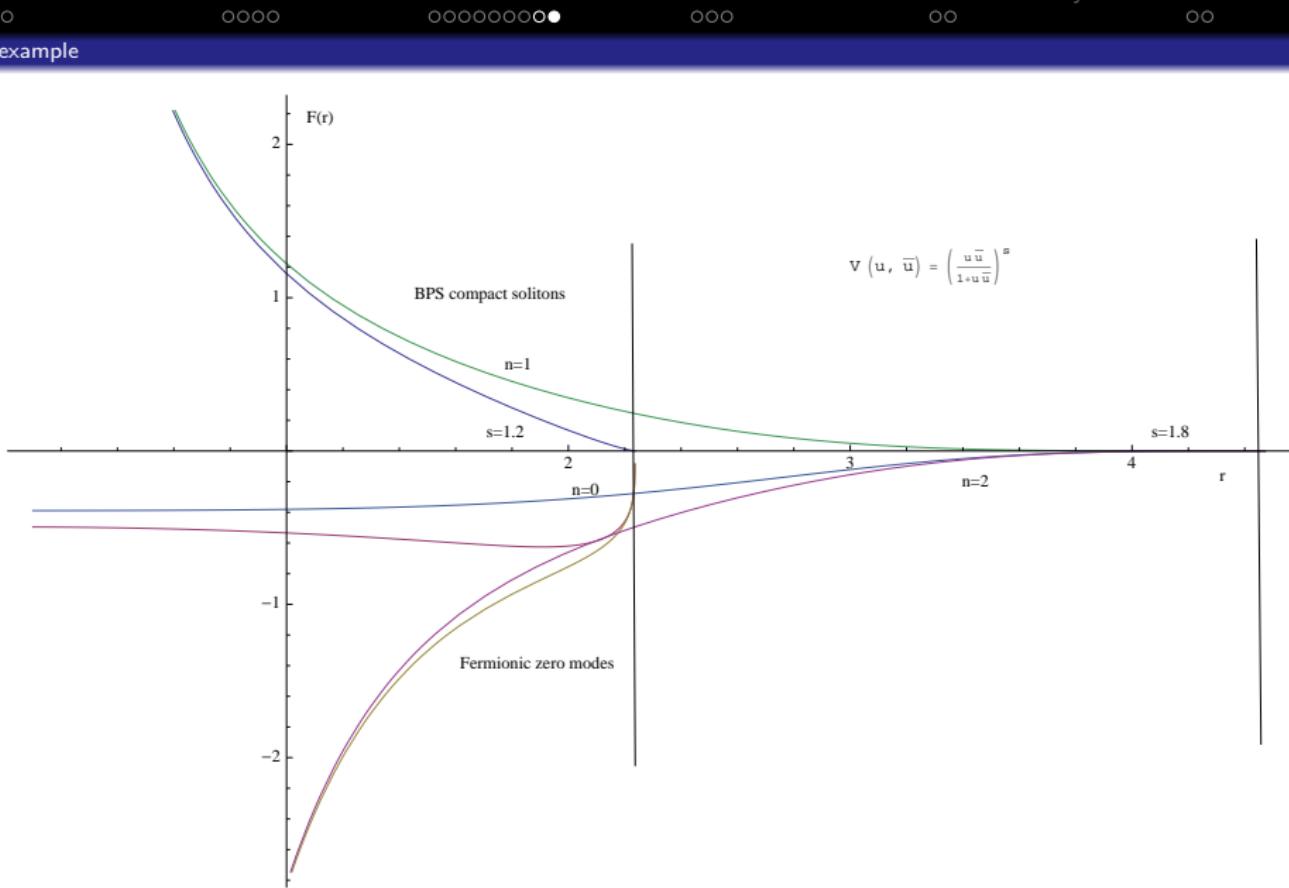


Figure: BPS compact solitons and fermionic zero modes

## Adding a D-term perturbation

- Non-compact BPS solutions  $\Rightarrow$  fermion zero modes are spread throughout space.
- Compact BPS solutions  $\Rightarrow$  fermion zero modes are confined to the compacton.

How general is this statement?

Let us consider the following action:

$$\begin{aligned}\mathcal{L} = & \int d^4\theta K(\Phi, \Phi^\dagger) + \int d^4\theta H(\Phi, \Phi^\dagger) D^\alpha \Phi D_\alpha \Phi \bar{D}^\dot{\beta} \Phi^\dagger \bar{D}_{\dot{\beta}} \Phi^\dagger \\ & + \left( \lambda \int d^2\theta W(\Phi) + \text{h.c.} \right)\end{aligned}$$

Expanding for small  $\lambda$

$$\begin{aligned}\mathcal{L}^{(1)} &= g(u, \bar{u}) u_{,\mu} \bar{u}^{\mu} + h(u, \bar{u})(u_{,\mu}^2)(\bar{u}_{,\nu})^2 + \mathcal{O}(\lambda^2) \\ \mathcal{L}^{(2)} &= h(u, \bar{u}) \left( (u_{,\mu})^2 (\bar{u}_{,\nu})^2 - (u_{,\mu} \bar{u}^{\mu})^2 \right) - \frac{g(u, \bar{u})^2}{4h(u, \bar{u})} \\ &\quad + 2\lambda |W'| \sqrt{-\frac{g(u, \bar{u})}{2h(u, \bar{u})} - u_{,\mu} \bar{u}^{\mu}} + \mathcal{O}(\lambda^2) \\ \mathcal{L}^{(3)} &= h(u, \bar{u}) \left( (u_{,\mu})^2 (\bar{u}_{,\nu})^2 - (u_{,\mu} \bar{u}^{\mu})^2 \right) - \frac{g(u, \bar{u})^2}{4h(u, \bar{u})} \\ &\quad - 2\lambda |W'| \sqrt{-\frac{g(u, \bar{u})}{2h(u, \bar{u})} - u_{,\mu} \bar{u}^{\mu}} + \mathcal{O}(\lambda^2)\end{aligned}$$

## Adding a D-term perturbation

Let us choose  $W'(\bar{u}) = m\bar{u}^l + \text{higher powers}$ . In a compacton configuration the profile function must behave close to the boundary as

$$f(r) = \beta \left( \frac{r - R}{R} \right)^\epsilon$$

Table: SUSY transformation at the boundary

	$\lim_{r \rightarrow R} \delta\psi^\alpha$
$l > \min(2 - \frac{2}{\epsilon}, s)$	0
$l < \min(2 - \frac{2}{\epsilon}, s)$	$\infty$
$l = \min(2 - \frac{2}{\epsilon}, s)$	$C\xi^\alpha$

In the third case the fermionic zero modes exist beyond the compacton support!

# SUSY baby Skyrme with quadratic term

We will use a trick (M. Nitta, S. Sasaki, Phys.Rev. D90 (2014) no.10, 105001) to construct the  $N = 2$  version of the full baby Skyrme model. SUSY Lagrangian again

$$\mathcal{L} = \int d^2\theta d^2\bar{\theta} K(\Phi, \Phi^\dagger) + \int d^2\theta d^2\bar{\theta} \Lambda D^\alpha \Phi D_\alpha \Phi \bar{D}^\beta \dot{\Phi}^\dagger \bar{D}_\beta \Phi^\dagger$$

where  $\Lambda$  is an unknown function of superfields and derivatives. Solving for  $F$

$$F = 0 \quad \text{and} \quad F\bar{F} = -\frac{g}{2\Lambda_0} - \partial_\mu u \partial^\mu \bar{u}$$

Substituting the trivial solution (this procedure works also for the non-trivial one)

$$\mathcal{L}|_{\text{on-shell}} = g \partial_\mu u \partial^\mu \bar{u} + \Lambda_0 (\partial_\mu u)^2 (\partial_\nu \bar{u})^2$$

Now force this Lagrangian to be the one corresponding to the full baby Skyrme model and solve for  $\Lambda_0$ .

$$\Lambda_0 = - \frac{V(u, \bar{u}) + h(u, \bar{u})(\partial_\mu u \partial^\mu \bar{u})^2 - h(u, \bar{u})(\partial_\mu u)^2(\partial_\nu \bar{u})^2}{(\partial_\mu u)^2(\partial_\nu \bar{u})^2}$$

promoting to superfields

$$\Lambda = - \frac{V(\Phi, \Phi^\dagger) + h(\Phi, \Phi^\dagger)(\partial_\mu \Phi \partial^\mu \Phi^\dagger)^2 - h(\Phi, \Phi^\dagger)(\partial_\mu \Phi)^2(\partial_\nu \Phi^\dagger)^2}{(\partial_\mu \Phi)^2(\partial_\nu \Phi^\dagger)^2}$$

(the term with four superderivatives saturates the Grassmannian integration for the bosonic sector). So, with this choice of  $\Lambda$

$$\mathcal{L} = \int d^2\theta d^2\bar{\theta} K(\Phi, \Phi^\dagger) + \int d^2\theta d^2\bar{\theta} \Lambda D^\alpha \Phi D_\alpha \Phi \bar{D}^{\dot{\beta}} \Phi^\dagger \bar{D}_{\dot{\beta}} \Phi^\dagger$$

generates the SUSY baby Skyrme model with quadratic term in the canonical branch ( $F = 0$ ).

# Conclusions

- We have studied the SUSY BPS baby Skyrme model.
- Generic BPS solutions preserve 1/4 SUSY.
- We have determined the fermion zero modes in the background of BPS solutions
  - Non-compact BPS solutions  $\Rightarrow$  fermion zero modes exist throughout space.
  - Compact BPS solutions  $\Rightarrow$  fermion zero modes are confined to the compacton support  $\Rightarrow$  restoration of SUSY outside the compacton.
- This behavior was observed for SUSY cosmic strings with higher derivative terms: Phys.Lett. B755 (2016) 498-503.
- The addition of a D-term allows for the "deconfinement" of fermion zero modes even in the background of a compact solution.

$\frac{1}{2}$  Thanks